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**Negligible Kinetic Helicity Can Drive Large Scale Dynamos**

JONATHAN PIETARILA GRAHAM, Los Alamos National Laboratory, ERIC BLACKMAN, University of Rochester, PABLO MININNI, NCAR & Universidad de Buenos Aires, ANNICK POUQUET, National Center for Atmospheric Research — Turbulent helical velocities drive large scale magnetic field growth and steepen the small scale magnetic energy spectrum, but the minimum sufficient fractional kinetic helicity  $f_{h,C}$  to do so has not been previously quantified. Using direct numerical simulations, we show that  $f_{h,C}$  strongly decreases as the ratio of forcing to large scale wavenumbers  $k_F/k_{min}$  increases. We also develop a simple theory that explains the simulation results. For  $k_F/k_{min} \geq 6$  we find  $f_{h,C} \lesssim 5\%$ , and our theory predicts that, in the asymptotic limit  $k_F/k_{min} \rightarrow \infty$ ,  $f_{h,C} \sim (k_F/k_{min})^{-5}$ , implying that very small helicity fractions strongly influence magnetic spectra for even moderate scale separation.

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