

Abstract Submitted
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Neoclassical Toroidal Viscosity Induced Rotation in Tokamaks and Quasi-symmetric Stellarators¹ A.J. COLE, C.C. HEGNA, J.D. CALLEN, University of Wisconsin — Non-axisymmetric magnetic perturbations generate variations in $|B|$ along a field line that induce non-ambipolar radial transport and a global toroidal force on the plasma, known as neoclassical toroidal viscosity [NTV]. A strong correlation exists between the flow evolution physics of tokamaks and quasi-helically symmetric [QHS] stellarators. In QHS-mode, there exists a helical symmetry angle $\alpha \equiv m\theta - n\zeta$, with m, n fixed integers that is analogous to the poloidal direction in tokamaks. As a result, there exists a direction of near helical symmetry and thus least flow damping along \vec{e}_h such that $\vec{e}_h \cdot \vec{\nabla}\alpha = 0$, analogous to the toroidal tokamak direction. In this paper, a model analytic 'toroidal' rotation equation is developed which smoothly transitions between previously asymptotic low-collisionality regimes [1], while incorporating both electron and ion NTV. In particular, the transition from ion to electron dominated NTV is presented in a single equation for the first time. This research will facilitate future comparison between NTV-induced rotation in QHS stellarators and tokamaks. [1] K.C. Shaing, Phys. Plasmas, **10**, 1443 (2003).

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