

Abstract Submitted
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A New Measurement of the Total Cross Section for the Photodisintegration of ${}^9\text{Be}$ Near Threshold¹ C.W. ARNOLD, T.B. CLEGG, H.J. KARWOWSKI, G.C. RICH, J.R. TOMPKINS, UNC Chapel Hill, TUNL, C.R. HOWELL, Duke University, TUNL — The hot and neutron rich dense matter in Type II supernovae is a plausible environment where r-process nuclei are synthesized. The $\alpha(\alpha n, \gamma){}^9\text{Be}(\alpha, n){}^{12}\text{C}$ sequence is the favored reaction chain for synthesizing carbon in this explosive environment. Nucleosynthesis network models indicate that the ratio of neutrons to seed nuclei at the onset of the r-process is highly sensitive to the rate of the $\alpha(\alpha n, \gamma){}^9\text{Be}(\alpha, n){}^{12}\text{C}$ chain relative to the $\alpha(\alpha\alpha, \gamma){}^{12}\text{C}$ reaction. The rate of the $\alpha(\alpha n, \gamma){}^9\text{Be}$ reaction is derived from the measured cross section for photodisintegration of ${}^9\text{Be}$. New cross section measurements for this reaction have been made from threshold to 5 MeV using the High Intensity Gamma-ray Source (HI γ S) at TUNL. The low energy spread (as low as 1%) of the beam at HI γ S enabled high precision measurements at the reaction threshold energy and of the narrow resonance at 2.43 MeV. Experimental techniques and results will be presented and astrophysical consequences will be discussed.

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