

Abstract for an Invited Paper  
for the DPP09 Meeting of  
The American Physical Society

### Gyrokinetic Simulations of Enhanced Alpha Transport by De-stabilized Alfvén Turbulence<sup>1</sup>

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Alfvén turbulence, destabilized by fusion-produced  $\alpha$ -particles, is expected to greatly enhance transport of these hot fusion products. Previously, the gyrokinetic code GYRO [1] was used to simulate the convective transport of fusion alpha particles by electrostatic ( $\beta = 0$ ) ITG/TEM turbulence driven at low  $k$  ( $0 < k_{\theta}\rho_s < 1$ ) by density and temperature gradients in the background plasma for the GA-standard case [2]. The present study, at  $\beta_e = 0.002$ , includes electromagnetic effects, allowing for driven Alfvén turbulence. The alpha particles are modeled by a hot Maxwellian ( $T_{\alpha} = 100 T_e$ ) superimposed at trace density ( $0.005 < n_{\alpha}/n_e < 0.025$  and  $a/L_{\alpha} = 4$  fixed) on the background plasma. Linear stability studies show two high-frequency modes driven unstable at very long wavelength ( $0 < k_{\theta}\rho_s < 0.14$ ) by a radial density gradient in the alpha population. A new eigenvalue solver within GYRO shows that the leading modes, identified as the toroidal Alfvén eigenmode (TAE) and the energetic particle mode (EPM), exhibit hybrid “drift-Alfvén” frequency scaling with  $k_{\theta}\rho_s$  and  $n_{\alpha}$ . At densities below the Alfvén linear stability threshold ( $n_{\alpha}/n_e \leq 0.005$ ), ITG/TEM turbulence dominates nonlinear simulations. In this limit, the transported alpha particles are passive tracers and ion and electron transport agrees well with  $n_{\alpha} = 0$  results. Just above the alpha-particle density gradient threshold, Alfvénic (TAE/EPM) drive enhances transport in  $\alpha$ -particle and background channels. This trend continues as  $n_{\alpha}$  and TAE/EPM microturbulence drive increase. The focus is on conditions for obtaining stationary nonlinearly saturated transport avoiding any subcritical limit on the total beta gradient.

[1] J. Candy and R.E. Waltz, Phys. Rev. Lett. **91**, 045001 (2003).

[2] C. Estrada-Mila, *et al.*, Phys. Plasmas **13**, 112303 (2006).

<sup>1</sup>Supported by US DOE under DE-FC02-08ER54977 and DE-FG03-95ER54309.