Abstract Submitted for the DPP10 Meeting of The American Physical Society

Linear and Nonlinear Whistler Wave Propagation in the Magnetosphere Based on Plasma Density Models from IMAGE Spacecraft **Data**<sup>1</sup> MAURICIO FLORES, Physics Department, The University of Texas at Brownsville, CYNTHIA CORREA, WENDEL HORTON, Institute for Fusion Studies, The University of Texas at Austin — From the radio plasma imager on the IM-AGE satellite, spatial profiles of electron density in the inner magnetosphere were constructed [B.W. Reinisch et al, Geophys. Res. Lett., 28, 1167 (2001)]. We use these profiles and the dipolar magnetic field model to analyze the propagation of whistler waves. We compute the dispersion characteristics of wave packets from the  $2D \ \omega(kx, kz, n, B)$  dispersion function, showing wave energy focusing into low phase velocity regions. We add model growth rates from S. Sazhin, Whistler-mode waves in a hot plasma (Cambridge U. Press, Cambridge, 1993) and nonlinear terms from Horton et al [W. Horton et al, Nonlinear Dynamics of the Electromagnetic Ion Cyclotron Structures, Firehose and Whistlers, preprint, Nonlin. Processes Geophys.] to determine saturation levels of whistler chorus waves and associated coherent structures. We explore NLS, DNLS and vortex models, consistent with experiments by Stenzel et al [R.L. Stenzel et al, Plasma Phys. Control. Fusion 50 074009 (2008)].

<sup>1</sup>Work partially supported by NSF Grant 0964692 to the University of Texas at Austin and by the OFES in the Dept of Energy.

Cynthia Correa Institute for Fusion Studies, The University of Texas at Austin

Date submitted: 18 Jul 2010

Electronic form version 1.4