Abstract Submitted for the DPP12 Meeting of The American Physical Society

Parametric dependence and inductive control of 3D helical equilibria in the MST RFP B.E. CHAPMAN, D.J. DEN HARTOG, S.T.A. KU-MAR, M. NORNBERG, E. PARKE, J.A. REUSCH, UW-Madison, S. CAPPELLO, P. FRANZ, P. PIOVESAN, M. PUIATTI, M. SPOLAORE, Consorzio RFX, D.L. BROWER, W.X. DING, L. LIN, UCLA — A stellarator-like equilibrium emerges in the core of RFP plasmas when the innermost resonant m = 1 tearing mode grows to large amplitude and the other, secondary m = 1 mode amplitudes are reduced. In MST, the likelihood and duration of these quasi-single-helicity (QSH) spectra increase strongly with Ip, similar to the trend observed in RFX-mod RFP plasmas which also develop a helical equilibrium. However, the Ip at which these spectra emerge in MST is nearly three times smaller than in RFX-mod. But due to additional differences in Te, Zeff, majority ion mass, and density, the two devices share a similar range of Lundquist number,  $S > 6 \ge 10^5$ , an important dimensionless parameter in resistive MHD, suggesting that S may be a predictive parameter for the onset of the helical equilibrium. At the largest S in MST, the amplitude of the dominant mode reaches 8% of the equilibrium field strength. The simultaneous reduction in secondary mode amplitudes leads to an increase in energy confinement time, reaching > 1 ms, about 50% larger than lower-S plasmas lacking a QSH spectrum. The secondary modes are reduced even further when the surface parallel electric field is inductively increased. This leads to a central Te > 1 keV and an energy confinement time  $\sim 3$  ms. There is also a 6 keV/m Te gradient in the core. Supported by USDOE.

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Date submitted: 18 Jul 2012

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