

Abstract Submitted
for the DPP17 Meeting of
The American Physical Society

Turbulent Dynamo Amplification of Magnetic Fields in Laser-Produced Plasmas: Simulations and Experiments¹ P TZEFERACOS, U Chicago, A RIGBY, A BOTT, A BELL, U Oxford, R BINGHAM, RAL, A CASNER, CEA, F CATTANEO, U Chicago, E CHURAZOV, MPIA Garching, C FOREST, U Wisconsin-Madison, J KATZ, LLE, M KOENIG, LULI, C-K LI, MIT, J MEINECKE, U Oxford, R PETRASSO, MIT, H-S PARK, B REMINGTON, J ROSS, D RYUTOV, LLNL, D RYU, UNIST, B REVILLE, Queens U Belfast, F MINIATI, ETH Zurich, A SCHEKOCHEV, U Oxford, D FROULA, LLE, D LAMB, U Chicago, G GREGORI, U Oxford — The universe is permeated by magnetic fields, with strengths ranging from a femtogauss in the voids between the filaments of galaxy clusters to several teragauss in black holes and neutron stars. The standard model for cosmological magnetic fields is the nonlinear amplification of seed fields via turbulent dynamo. We have conceived experiments to demonstrate and study the turbulent dynamo mechanism in the laboratory. Here, we describe the design of these experiments through large-scale 3D FLASH simulations on the Mira supercomputer at ANL, and the laser-driven experiments we conducted with the OMEGA laser at LLE. Our results indicate that turbulence is capable of rapidly amplifying seed fields to near equipartition with the turbulent fluid motions.

¹This work was supported in part from the ERC (FP7/2007-2013, no. 256973 and 247039), and the U.S. DOE, Contract no. B591485 to LLNL, FWP 57789 to ANL, grant no. DE-NA0002724 and DE-SC0016566 to the University of Chicago, and DE-AC02-06CH11357 to ANL.

Petros Tzeferacos
U Chicago

Date submitted: 14 Jul 2017

Electronic form version 1.4