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Extracting ANCs in neutron transfer reactions to determine proton capture reaction rates TARIQ AL-ALBDULLAH, X. CHEN, C.A. GAGLIARDI, Y.-W. LUI, G. TABACARU, Y. TOKIMOTO, L. TRACHE, R.E. TRIBBLE, Texas A&M University, F. CARSTOIU, Institute of Physics and Nuclear Engineering, Romanai — The high temperatures (>10⁸ K) in novae outbursts enable unstable nuclei to leak out from the hot CNO cycle to the rp-process, where heavier nuclei such as ¹⁸F and ²²Na are synthesized and might be ejected. Their abundances can be influenced by the ${}^{17}{\rm F}({\rm p},\gamma){}^{18}{\rm Ne}$ and ${}^{22}{\rm Mg}({\rm p},\gamma){}^{23}{\rm Al}$ reactions respectively. The first reaction connects the CNO and NeNa cycles, while the second may explain the unobserved γ -ray emission from ²²Ne due to the β decay in ²²Na. We have applied an indirect technique to determine the above reaction rates at stellar energies. We have measured the neutron transfer reactions ¹³C(¹⁷O, ¹⁸O)¹²C and ¹³C(²²Ne, ²³Ne)¹²C to determine the asymptotic normalization coefficients (ANCs) for the ground and first excited states in ¹⁸O and ²³Ne. These ANCs can be transposed to the corresponding states in the mirror nuclei ¹⁸Ne and ²³Al respectively. As a part of these experiments, we have measured the elastic scattering data to obtain the optical model parameters that are used in DWBA calculations, and hence to extract the ANCs.

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