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Photoactivation of ^{176m}Lu via Bremsstrahlung at the Stuttgart DYNAMITRON BRIAN GODDARD, Drexel University, TOM HENRY, University of Surrey, TREVOR BALINT, Youngstown State University, HEINZ-HERMANN PITZ, FRANK STEDILE, ULRICH KNEISSL, University of Stuttgart, JEREMY GAISON, Drexel University, CHRIS CHIARA, US Army Research Laboratory, TRISTAN WINICK, Drexel University, JEFF CARROLL, US Army Research Laboratory — The ground state of ¹⁷⁶Lu has a very long half-life of approximately 37 billion years and primarily β^- decays (>99.9%) to ¹⁷⁶Hf. However, ¹⁷⁶Lu possesses an isomer ($J^{\pi} = 7^{-}$) 123 keV above the ground state ($J^{\pi} = 1^{-}$) that also β^- decays to $^{176}{\rm Hf}$ but with a much shorter half-life of about 3.6 hours. The study of this isomer could lead to new findings regarding astrophysical nucleosynthesis. A disparity between the predicted abundance of ¹⁷⁶Lu due to nucleosynthesis and the actual measured abundance implies that transitions from the isomer to the ground state via intermediate states must have taken place during the s-process. Since the rates at which these transitions occur are temperature dependent, ¹⁷⁶Lu could be used as an s-process "thermometer." A photoactivation experiment was performed on ¹⁷⁶Lu at the Stuttgart DYNAMITRON using bremsstrahlung with varying endpoints between 0.7 and 2.2 MeV to determine the intermediate state energies and integral cross sections for the transitions that lead to the isomer. We present the results of the analysis of the data as well as preliminary values for the intermediate state energies and their cross sections.

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