

Abstract Submitted  
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**Measurement of the Total Cross-Section for the  ${}^9\text{Be}(\gamma, n\alpha)\alpha$  reaction**<sup>1</sup> C.W. ARNOLD, T.B. CLEGG, H.J. KARWOWSKI, UNC-Chapel Hill, TUNL, C.R. HOWELL, A.P. TONCHEV, G. RUSEV, Duke University, TUNL — The  ${}^9\text{Be}(\gamma, n\alpha)\alpha$  cross section is key to understanding isotopic abundances of nuclei produced during the r-process. The inverse reaction bridges the unstable mass gaps at A=5 and 8 leading to  $\alpha(\alpha n, \gamma){}^9\text{Be}(\alpha, n){}^{12}\text{C}$  and so on, producing seed nuclei for the r-process and setting the neutron-to-seed nucleus ratio that drives universal isotopic abundance predictions [Ref 1,2]. In order to make high precision measurements ( $\pm 5\%$ ) of the  ${}^9\text{Be}(\gamma, n\alpha)\alpha$  cross-section which includes narrow resonances, tunable gamma ray beam with small  $\Delta E/E$  is required along with gamma and neutron detectors whose efficiencies are well known. We used TUNL's high intensity gamma ray source (HI $\gamma$ S) to measure the cross sections for  ${}^9\text{Be}(\gamma, n)$  in the energy range of 1.55 to 5.0 MeV with beam energy resolutions between 14 and 150 keV as determined by large Ge detector. The neutrons were detected using  ${}^3\text{He}$  proportional counter. Experimental details will be discussed and the results as well as their astrophysical consequences will be presented. [Ref 1] B. Meyer *et al.*, *Astro J.*, **399** 656-664 (1992). [Ref 2] T. Kajino *et al.*, *Nuc. Phys. A*, **704**, 165c-178c (2002)

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