

SES14-2014-000085

Abstract for an Invited Paper  
for the SES14 Meeting of  
the American Physical Society

**Full-dimensional quantum dynamics calculations of rovibrationally inelastic scattering of CO-H<sub>2</sub><sup>1</sup>**

BENHUI YANG, University of Georgia

The CO-H<sub>2</sub> collisional system is crucial in determining the physics and chemistry of interstellar environments due to the astrophysical importance of H<sub>2</sub> and CO. We calculated for the first time a full-dimensional (6D) potential energy surface (PES) for this system using the high-level CCSD(T)-F12B method. The PES was fitted using an invariant polynomial method in 6D. Quantum close-coupling calculations of rotational and vibrational quenching of CO in collisions with H<sub>2</sub> were carried out on the new 6D PES. The pure state-to-state rotational excitations from CO( $v_1 = 0, j_1=0, 1$ ) were benchmarked with crossed molecular beam measurement for collision energies of 795 - 991 cm<sup>-1</sup> and 3.3 - 22.5 cm<sup>-1</sup>. The computed cross sections for  $j_1 = 0 \rightarrow 1$  transition in CO and show better agreement with measurement than those obtained on a recently available 4D PES. For rovibrational transitions, state-to-state and total quenching cross sections and rate coefficients were calculated for the vibrational quenching in CO( $v_1 = 1, j_1$ )+H<sub>2</sub>( $v_2 = 0, j_2$ )  $\rightarrow$  CO( $v'_1 = 0, j'_1$ )+H<sub>2</sub>( $v'_2 = 0, j'_2$ ) collisions,  $j_1 = 0, 2$  for para-H<sub>2</sub> and  $j_1 = 1, 3$  for ortho-H<sub>2</sub>. The results are compared with experimental results and previous calculations using 4D PESs and various decoupling approximations. Our calculation also confirmed that the contribution from a quasi-resonant channel, CO( $v_1=1$ ) + H<sub>2</sub>( $v_2=0, j_2=2$ )  $\rightarrow$  CO( $v'_1=0$ ) + H<sub>2</sub>( $v'_2 = 0, j'_2=6$ ), dominates the vibrational quenching of CO in collision with para-H<sub>2</sub> for  $T \geq 50$  K.

<sup>1</sup>Work at UGA and Emory are supported by NASA grant No. NNX12AF42G, at UNLV by NSF Grant No. PHY-1205838, and at Penn State by NSF Grant No. PHY-1203228.