

Abstract Submitted
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Magnetohydrodynamic Bondi–Hoyle Accretion¹ ANDREW PENNER, University of British Columbia — In relativistic Bondi–Hoyle accretion, a massive body travels through a uniform space background. As this body accretes matter it is known to develop either a bow shock or a tail shock. For the parameter space where a bow shock develops, there is the potential for a flip-flop instability to develop. Previous studies of the relativistic extension by Font et al. revealed that the flip-flop instability did not develop; however, they caution that their parameter space, when taken in the Newtonian limit, was outside of the regime that a Newtonian flip-flop would occur. In my research, I extend the works of Font et al. to include the effects of ideal magnetohydrodynamics. First, I continue the search for the flip-flop instability by analyzing lower asymptotic speeds of sound, which correspond closer to the Newtonian region in which a flip-flop will occur. I then include the magnetic fields to determine if these extra fields can trigger further instabilities and possibly enhance the likelihood of developing a flip-flop instability.

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