

Abstract Submitted  
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**Alfven Wave Propagation in Young Stellar Systems** RAY HUM-  
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systems have disks that are threaded by magnetic field lines with an hourglass ge-  
ometry. These fields funnel ionizing cosmic rays (CRs) into the system. However,  
the effect is offset by magnetic mirroring. An previous analysis considered how the  
presence of magnetic turbulence moving outward from the disk would effect the  
propagation of cosmic-rays, and in turn, change the cosmic-ray ionization fraction  
occurring within the disk. This work indicated that turbulence reduces the overall  
flux of cosmic-rays at the disk, which has important consequences for both chemical  
processes and planet formation that occur within these environments. However, the  
analysis assumed ideal MHD condition in which the gas is perfectly coupled to the  
magnetic field. We explore here the validity of this assumption by solving the full  
equations governing the motion of both ions and neutral within the system.

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